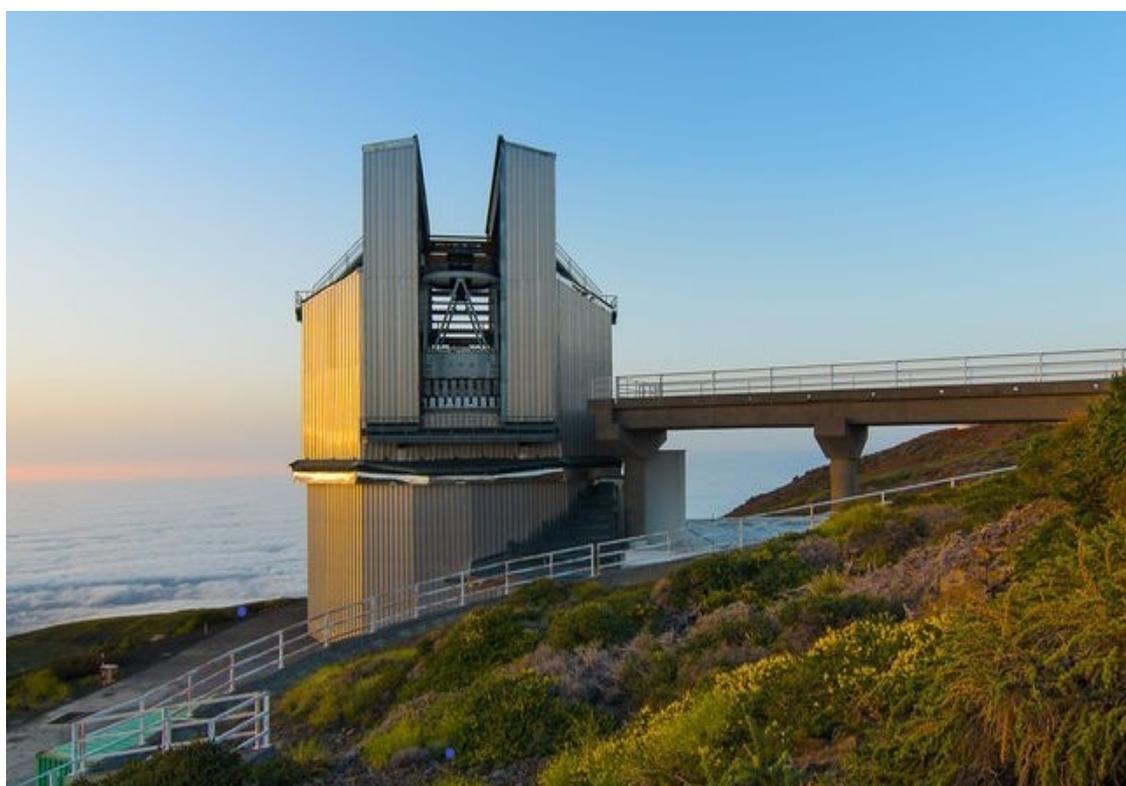


Observing at TNG today. Perspectives for the future.

Ennio Poretti



Instrument	Date
TNG	June, 9 th 1998
OIG	Dec, 10 th 1998
ARNICA	Dec, 18 th 1998
AdOpt	Dec, 18 th 1998
DOLORES	May, 20 th 2000
SARG	June, 9 th 2000
NICS	September, 17 th 2000
HARPS-N	March, 21 st 2012
GIANO	July, 27 th 2012
GIANO-B	Oct, 27 th 2016
GIARPS	March, 14 th 2017

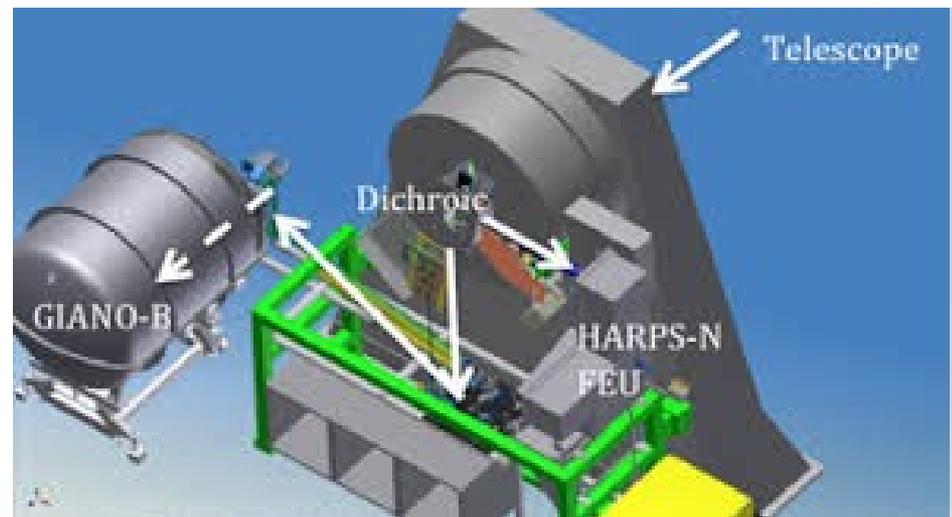
M1 diameter	3.58 m
Focal length	38.5 m (f/11)
M2 diameter	0.875 m
M2 baffle diam.	1.165 m
Scale	5.36 arcsec/mm
Vignetting-free field	25 arcmin diameter

Four interchangeable instruments

- **HARPS-N**, high-resolution spectrograph ($R=115000$) operating in the visible.
 - **GIANO-B**, high-resolution spectrograph ($R=50000$) operating in the near infrared.
 - **DOLORES**, low-resolution spectrograph ($R<6000$) and imaging.
 - **NICS**, near-infrared instrument, low-resolution spectroscopy ($R<2500$), imaging.
- and**
- **SiFAP2**, ultrarapid photometer (time resolution 8 ns), PI instrument.

GIANO-B and HARPS-N combined in the **GIARPS** observing mode. Simultaneous visible and infrared spectra of the same target.

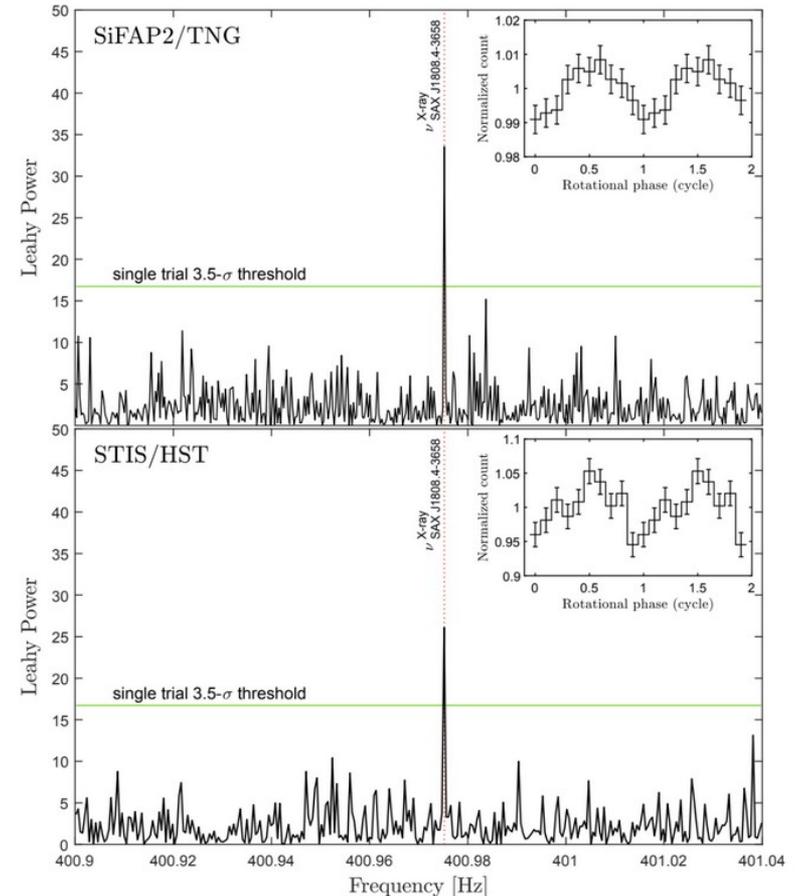
Action taken by the Italian Community : GAPS, GIANO Team, and TNG staff. Funded by own budget.



ENTERING TIME DOMAIN ASTRONOMY WITH EXCELLENCE

Silicon Fast optical Astronomical Photometer (SiFAP2) offered as a visitor instrument

**Pulses of visible light from
a millisecond pulsar**
(Ambrosino et al. 2017, *Nature Astronomy*)



First light of SiFAP2 on November 14, 2018

Ambrosino et al. 2020, Nature Astronomy

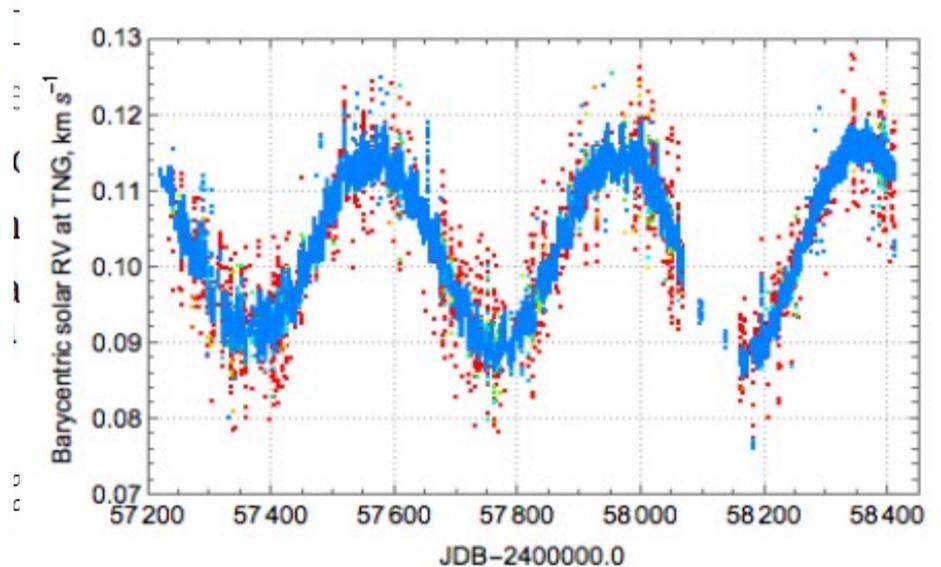
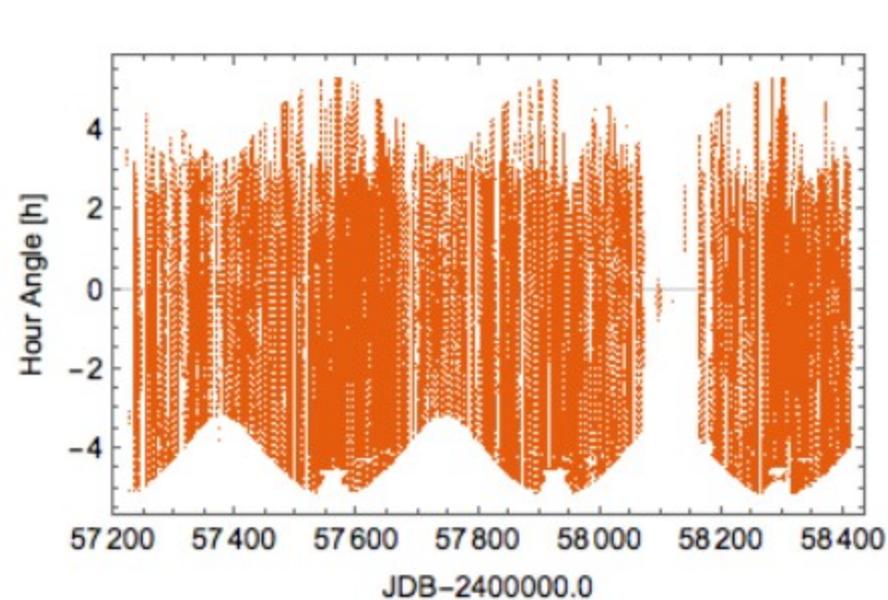
The Sun as a star

David Phillips, Xavier Dumusque, TNG staff, et al.

LCST (Low Cost Solar Telescope) operating daytime.
It feeds HARPS-N spectrograph.



Three years of Sun-as-a-star radial-velocity observations on the approach to solar minimum.



LOCNES: Low Cost NIR Extended Solar Telescope

Claudi R.^a, Ghedina A.^b, Pace E.^c, Gallorini L.^c, Di Giorgio A.-M.^d, Liu S.-J.^d, Tozzi A.^e, Carleo I.^a, Lanza A.F.^f, Micela G.^g, Molinari E.^h, Poretti E.^b, Phillips D.^g, and Tripodo G.ⁱ

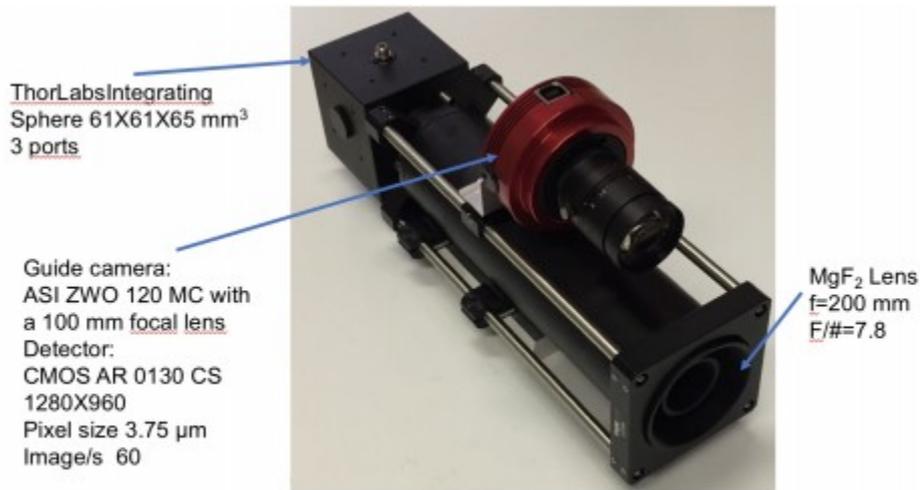
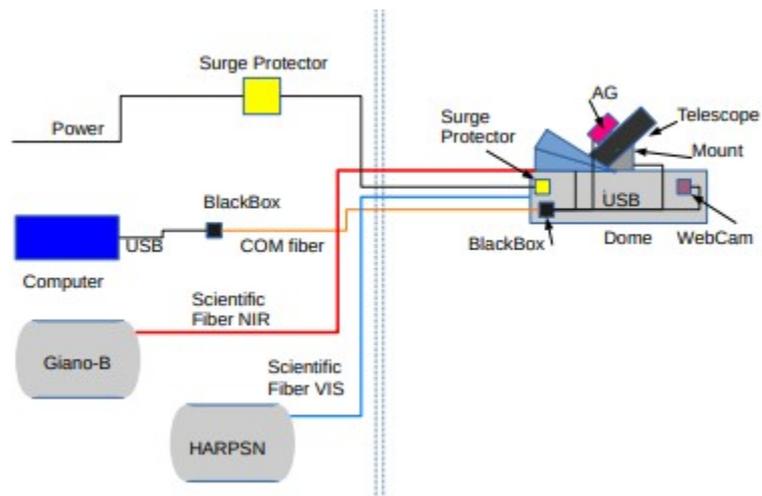


Figure 3. The LOCNES telescope.



The Fundación Galileo Galilei-INAF (FGG-INAF) is a Spanish non-profit institution. The FGG's main aim is to promote astrophysical research by managing the Telescopio Nazionale Galileo (TNG), located at the Roque de Los Muchachos Observatory in La Palma, and developing INAF activities in the Canary Islands.

The financial support is ensured by the Italian Istituto Nazionale di Astrofisica (INAF). Decisional Board : **PATRONATO**, composed of 5 members: INAF President, INAF General Director, INAF Scientific Director, and two experts.

On the island: TNG Director. Also acting as *Gerente* FGG.
31 personnel units contracted by FGG.

4 INAF astronomers.

Administration (1+2), Technology (11), Informatics (4), Astronomy (8), Telescope Operators (5), Safety Manager (1), **ASTRI (2)**

QUESTIONS

- Q1 HOW ARE NIGHTS AT TNG ?
- Q2 WHO USES TNG TIME ?
- Q3 IS IT DIFFICULT TO GET OBS.TIME ?
- Q4 CAN I GET ENOUGH DATA ?
- Q5 WHAT DOES TNG OBSERVE ?
- Q6 ARE THE RESULTS PUBLISHED?
- Q7 WHO PUBLISHES BY USING HR ?

Q1 : HOW ARE THE NIGHTS AT TNG ?

Night Statistics

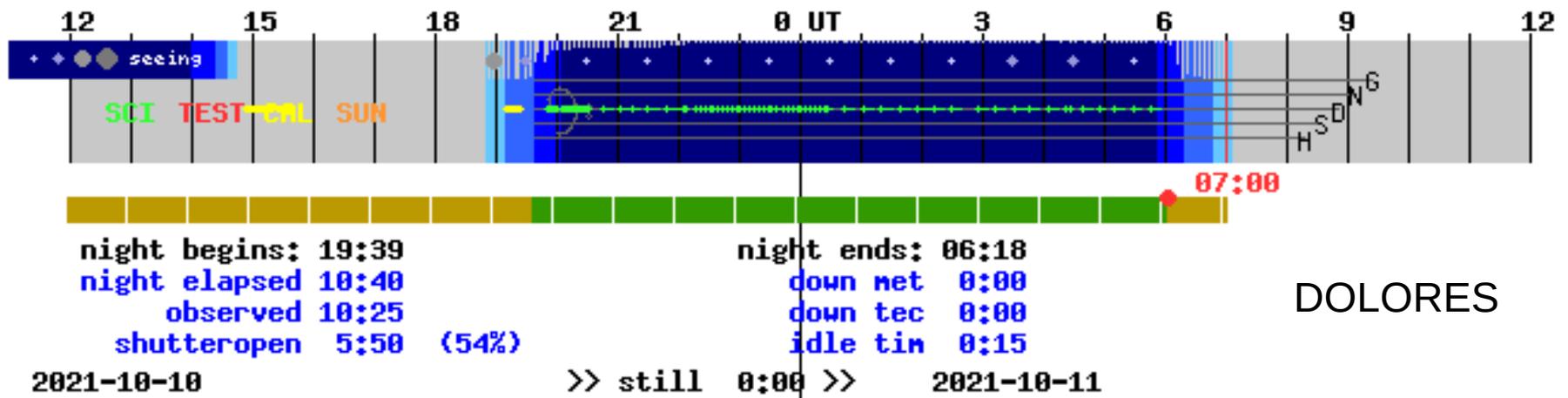
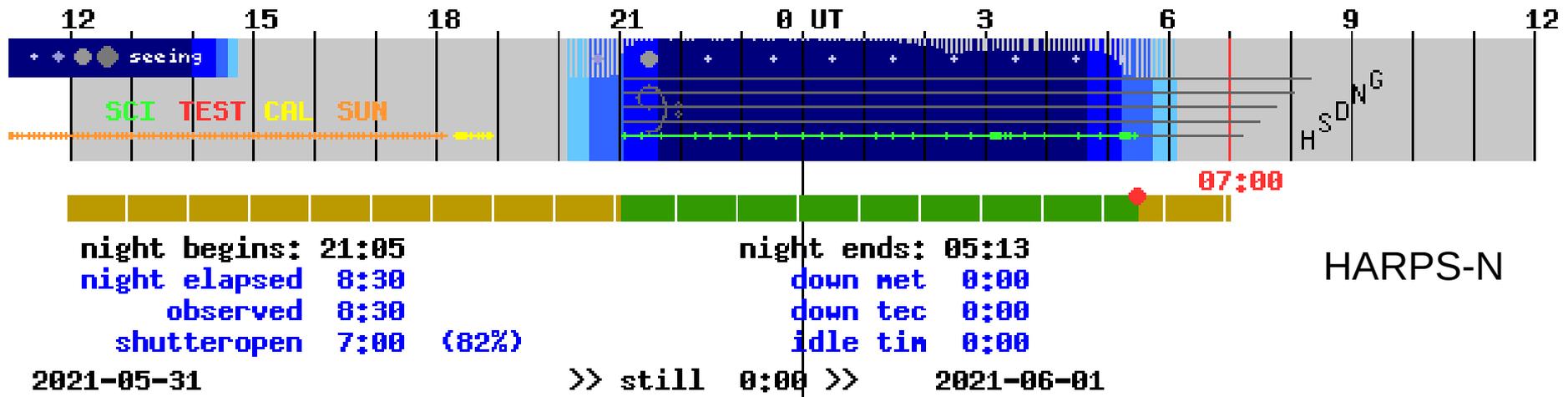
	2020	2019	2018	2017	2016	2015
	[%]	[%]	[%]	[%]	[%]	[%]
Down meteo	26.5	14.4	27.0	22.7	25.3	23.3
Technical failures	2.7	3.1	2.8	2.0	1.4	2.2
Engineering time	2.2	5.9	5.3	5.6	4.9	2.7
Idle time	0.5	0.6	0.5	0.6	1.0	0.4
Observed time	68.1	76.0	64.6	68.9	67.3	71.4
Open Shutter	74.6	71.9	76.2	72.8	74.1	72.7

Answers :

2.5 nights out of 10 lost for bad weather

You observe 7 nights out of 10

Open shutter 5.25 nights out of 10



Q2: WHO USES TNG TIME ?	AOT 38	AOT 39	AOT 40	AOT 41	AOT 42	AOT 43	AOT 44	Typical Semester
								NIGHTS
INAF OPEN TIME	12	12	10	18	16	26	28	20
INAF Large Prog.	47	47	39	43	43	34	36	40
INAF Long Term	-	4	3	-	6	6	-	3
GTO (INAF-GTO agr.)	40	40	33	40	40	40	40	40
Spanish CAT (Intl. protocol)	31	31	26	31	31	31	31	31
CCI ITP (Intl. protocol)	8	7	7	8	7	7	8	8
OPTICON	10	10	10	10	10	10	10	10
TNG-NOT (Agreement)	10	10	8	5	5	5	5	5
Technical nights (DDT, payback,...)	24	22	42 Alum.	28 PAO	24	24	24	24

ANSWERS :

63 nights INAF

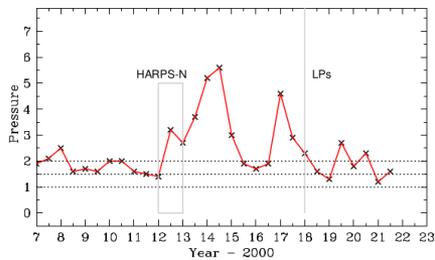
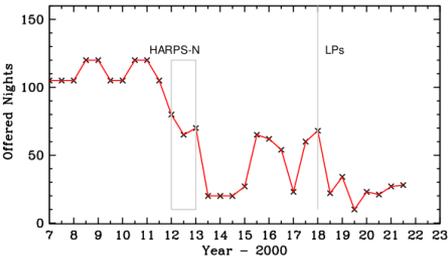
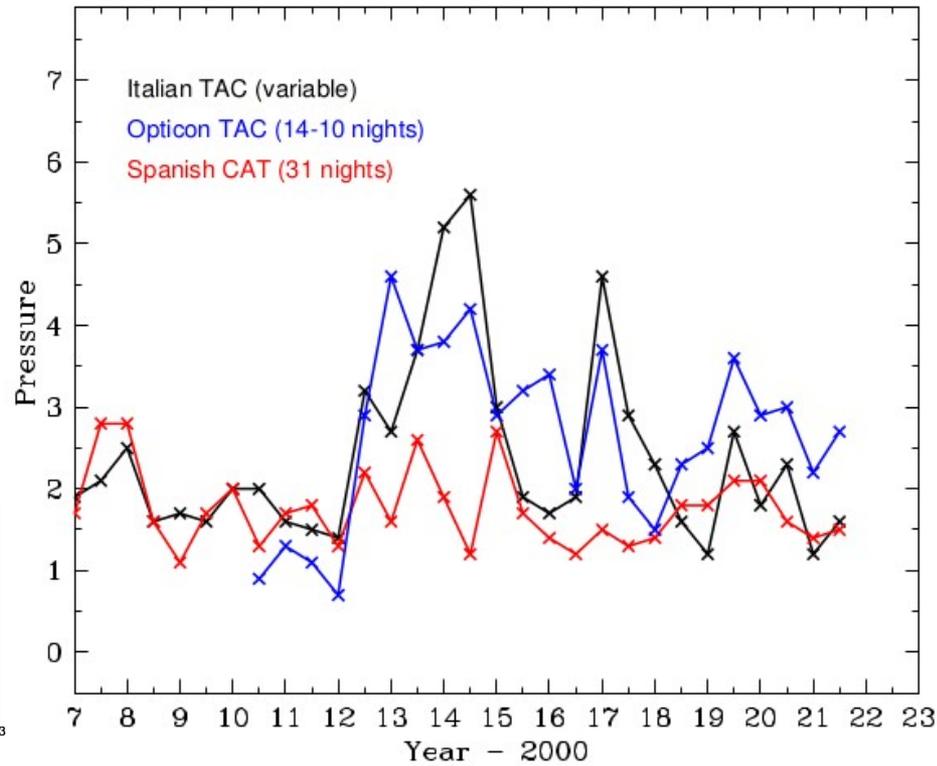
40 nights GTO

31 nights Spain

23 nights Rest-of-world

24 nights "Technical"

Q3 : IS IT DIFFICULT TO GET OBSERVING TIME ? HOW MUCH PRESSURE ?



Answer : Enough pressure to guarantee “good science”

Q4: CAN I GET ENOUGH DAT ?

According to the official table with the status of SPA observations made available to us by the TNG Director, 596 hrs of open telescope was recorded out to the 612 hrs nominally scheduled.

During the 596 hrs of observations in the six semesters from Period 37 to Period 42, we secured spectra for 520 stars in the disk field and associated star clusters, spreading a wide range of Galactocentric distances and divided in three main age groups, as shown in Figure 1 and detailed below .

Other cases : programs requesting <20-30 hours of observations mostly fully-executed
MINCE (Cescutti), C-MetaLL (Ripepi), WD (Silvotti), TESS subgiants (Montalto),
ARIEL (Rainer, Benatti)

More difficult to be fully executed: comets, asteroids and NEO due to high proper motion

Flexible schedule.

ANSWER : Yes, for sure

Q5: WHAT DOES TNG OBSERVE ?

OBSERVING PROGRAMS: INAF (63 nights/semester)



HIGH-RESOLUTION (GIARPS)

EXOPLANETS : GAPS and a few related PIs

STELLAR ASTROPHYSICS : ABUNDANCES MW DISK (SPA), METAL POOR (MINCE, Cescutti), T TAURI (GHOST Antonucci, JEDI Alcalá), GAIA CEPHEIDS (Ripepi), WD (Silvotti)

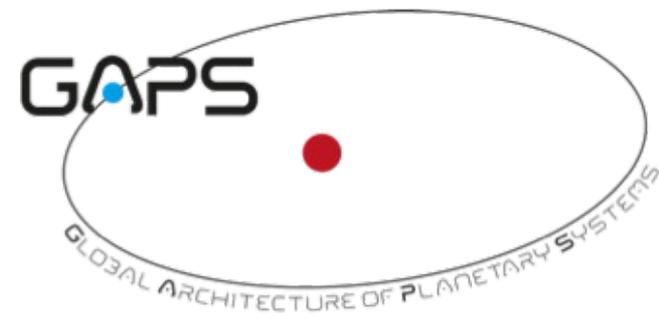
LOW RESOLUTION, IMAGING (DOLORES, NICS)

EXTRAGALACTIC : GW (GRAWITA), GRB (Swift), QSO (Trieste), SN (Padova), GALAXY CLUSTERS

SOLAR SYSTEMS : COMETS (Lazzarin), NEO (Perna), ASTEROIDS (IAPS, ESA/NASA preparatory work)

TIME DOMAIN ASTRONOMY (SiFAP2)

PULSARS (Ambrosino, Papitto, Casella,...), BURSTS, LUNAR AND ASTEROIDAL OCCULTATIONS



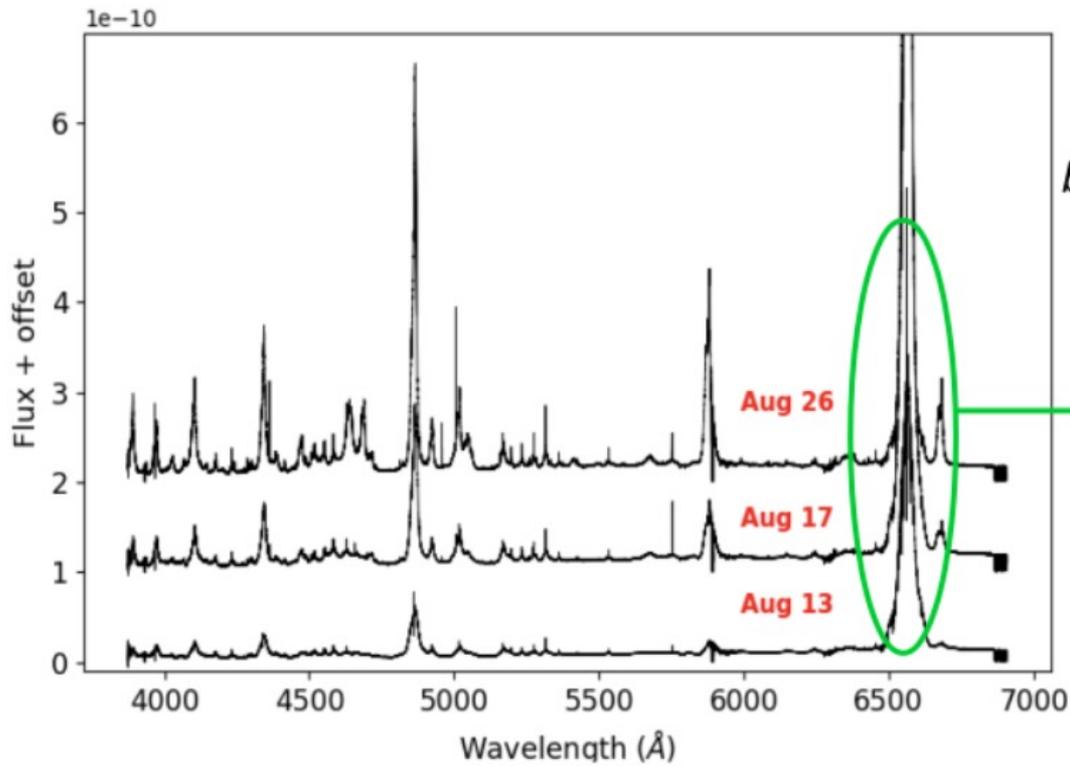
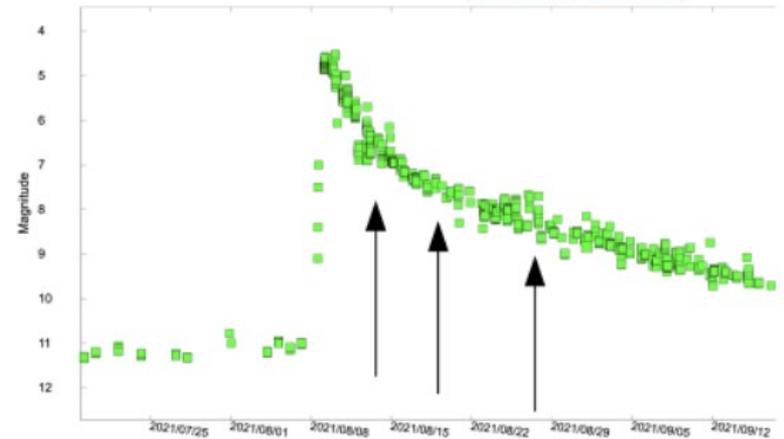
RS Oph & HARPS-N

(Luca Izzo)

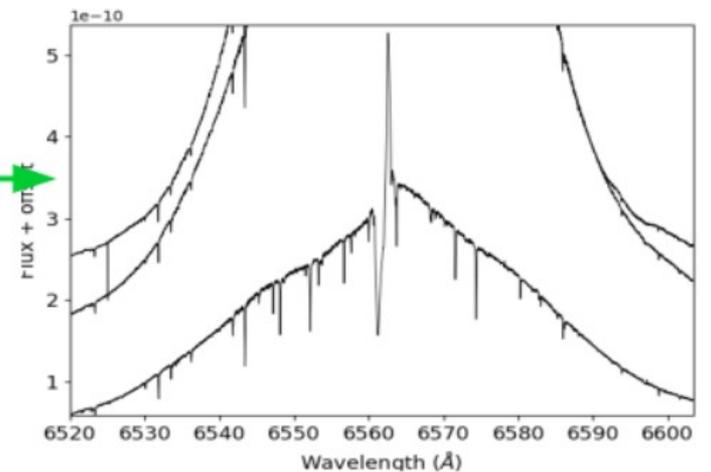
Recurrent nova – last outburst 2006

3 spectra obtained in August 2021

Within a larger high-res monitoring program:
UVES+TNG+NOT



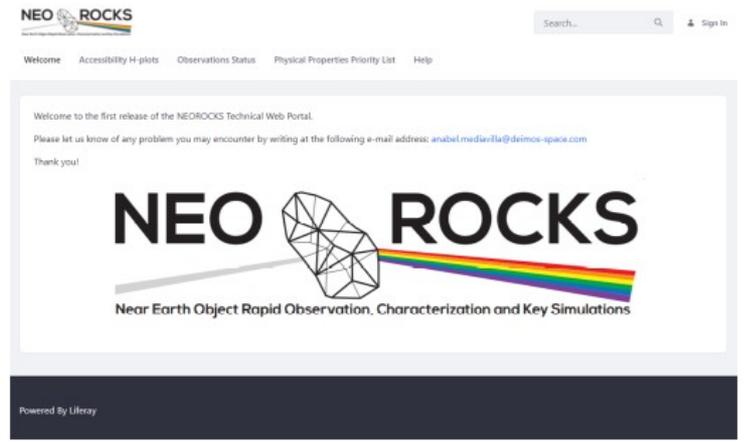
Shock interaction
nova ejecta – red giant wind
broad -> nova ejecta; *narrow* -> RG wind



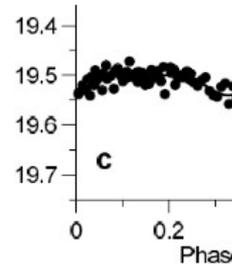
International commitments:



EU H2020 projects



Space missions



About 50 papers in 15 years: observations, interpretation, models, “big picture”

OBSERVING PROGRAMS: SPAIN (31 nights/semester)

HIGH-RESOLUTION (GIARPS)

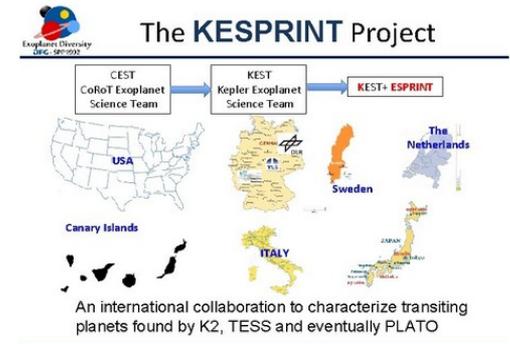
EXOPLANETS : MS STARS (IAC, KESPRINT), TRANSMISSION SPECTROSCOPY (IAC, Palles'),
M-STARS (IAC Rebolo, HADES, GAPS coll.)

STELLAR ASTROPHYSICS & EXOPLANETS: TAPAS (CSIC)

LOW RESOLUTION, IMAGING (DOLORES, NICS)

EXTRAGALACTIC : BLAZARS (Madrid), GALAXY CLUSTERS (IAC)

SOLAR SYSTEMS : –



OBSERVING PROGRAMS: INTERNATIONAL (CCI, OPTICON, and now NORDIC time)

HIGH-RESOLUTION (GIARPS)

EXOPLANETS : TRANSITING, TRANSMISSION SPECTROSCOPY

STELLAR ASTROPHYSICS: T TAURI, RCB

LOW RESOLUTION, IMAGING (DOLORES, NICS)

EXTRAGALACTIC : SN (ZTF, Icecube, Lensed...), AGN (eROSITA)

SOLAR SYSTEMS : COMETS, ASTEROIDS



PARTIAL RENEWAL OF THE BILATERAL AGREEMENT IN 2022

(CCI President Welcome speech)

Q5 WHAT DOES TNG OBSERVE ?

Answers:

Exoplanets (masses, transmission spectroscopy) by different groups, with a large INAF involvement.

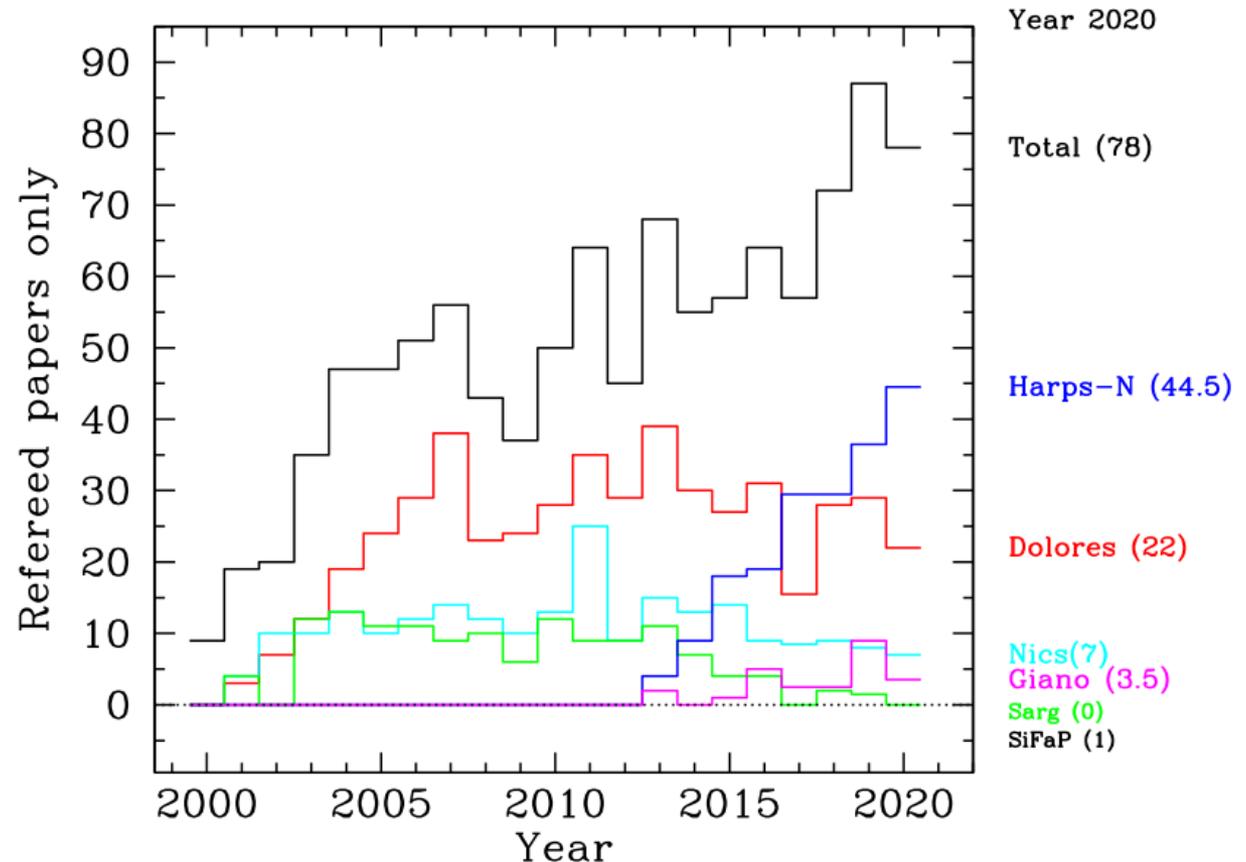
Stellar astrophysics, Solar System, TDA and extragalactic researches are adequately covered.

High-resolution spectroscopy. But also imaging, LR spectroscopy, fast photometry are powerful available tools.

Q6 : ARE THE RESULTS PUBLISHED ?

Publication record *(source TNG webpage)*

Peer-review journals only



Answers : Increasing after a long-standstill.

Large-Program (GAPS,GTO,SPA) effect since Dolores/Nics stable over the years

Q7: WHO PUBLISHES BY USING HR ?

Publication record (source TNG webpage)

Peer-review journals only

	2013	2014	2015	2016	2017	2018	2019	2020	
HARPS-N	4	9	22.5	22	31	29	33	40	
High Resolution	6	9	24.5	27	35	32	45	48	
INAF	4	2	11	9.5	9	9	14	17.5	
GTO	1	3	5	6	6	5	11	9	
SPAIN	0	1	3.5	7.5	8	9	10	12.5	
Rest of World (archival data!)	1	3	5	4	12	9	10	9	

ANSWER : INAF HOLDS THE LEAD

Q & A

- | | | |
|----|-----------------------------------|--|
| Q1 | HOW ARE NIGHTS AT TNG ? | It sounds reasonable to apply |
| Q2 | WHO USES TNG TIME ? | Italy, but not only |
| Q3 | IS IT DIFFICULT TO GET OBS.TIME ? | No, easy for good science |
| Q4 | CAN I GET ENOUGH DATA ? | For sure |
| Q5 | WHAT DOES TNG OBSERVE ? | Mostly Exoplanets in HR mode, but not only |
| Q6 | ARE THE RESULTS PUBLISHED? | Yes, ever increasing |
| Q7 | WHO PUBLISHES BY USING HR ? | INAF holds the lead |

Padova meeting on TNG
HARPS-N agreement renewed
Call for Large Programs (80-85% of available nights)
Two LPs using HR selected

March 2017
August 2017
October 2017
March 2018

This strategy is keeping TNG at the top in the HR field.

Date : December 7, 2018



Collaboration Agreement

for the maintenance and operation of the HARPS-N
spectrograph and its scientific use



CRITICALITIES AND FUTURE ACTIVITIES

HARPS-N: excellence in the exoplanet science in the Northern Hemisphere. Space missions on exoplanets (TESS, CHEOPS, ARIEL, PLATO,...) ensure a bright future.



GIANO-B: infrared spectrograph to characterize stars and exoplanets.

DOLORES and NICS are playing the game of transients follow-up and chasing the optical counterpart of a gravitational wave event. Small improvements possible

New instruments: SiFAP2, LOCNES, waiting for a third generation?

GIANO-B : Unable to reach high-precision RV measurements.
Instrument instability. Few lines in the IR.
Possible solution to alleviate the problem:
absorption cells

